

Development of Superconducting Tunnel Junction Detector using Hafnium (Hf-STJ)

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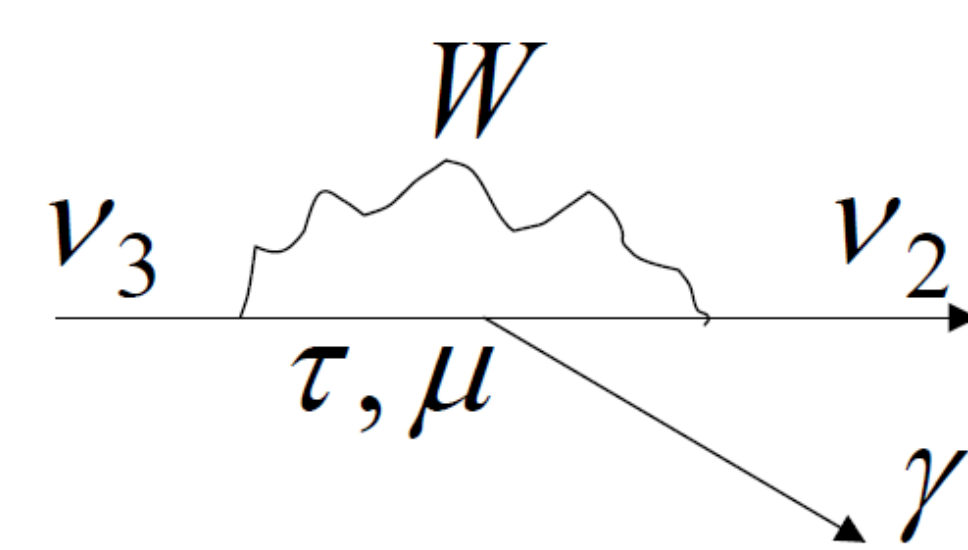
1. Introduction

- Difference between mass-square of different generation neutrino has been measured by various experiments of neutrino oscillation.
- However, neutrino mass itself has not been measured.
- Detection of neutrino decay enable us to measure an independent quantity of the difference between squares of neutrino mass.
- **→ We can obtain neutrino mass itself** from these two independent measurements.
- To observe neutrino decay, we need measure the energy spectrum of the cosmic infrared photon background with energy resolution better than 2%.
- We adopted the Hf-STJ detector as such a detector and are developing it.

2. Neutrino Decay

- In neutrino decay process, a lighter neutrino(ν_2) and a photon(γ) are emitted from a heavier neutrino(ν_3).

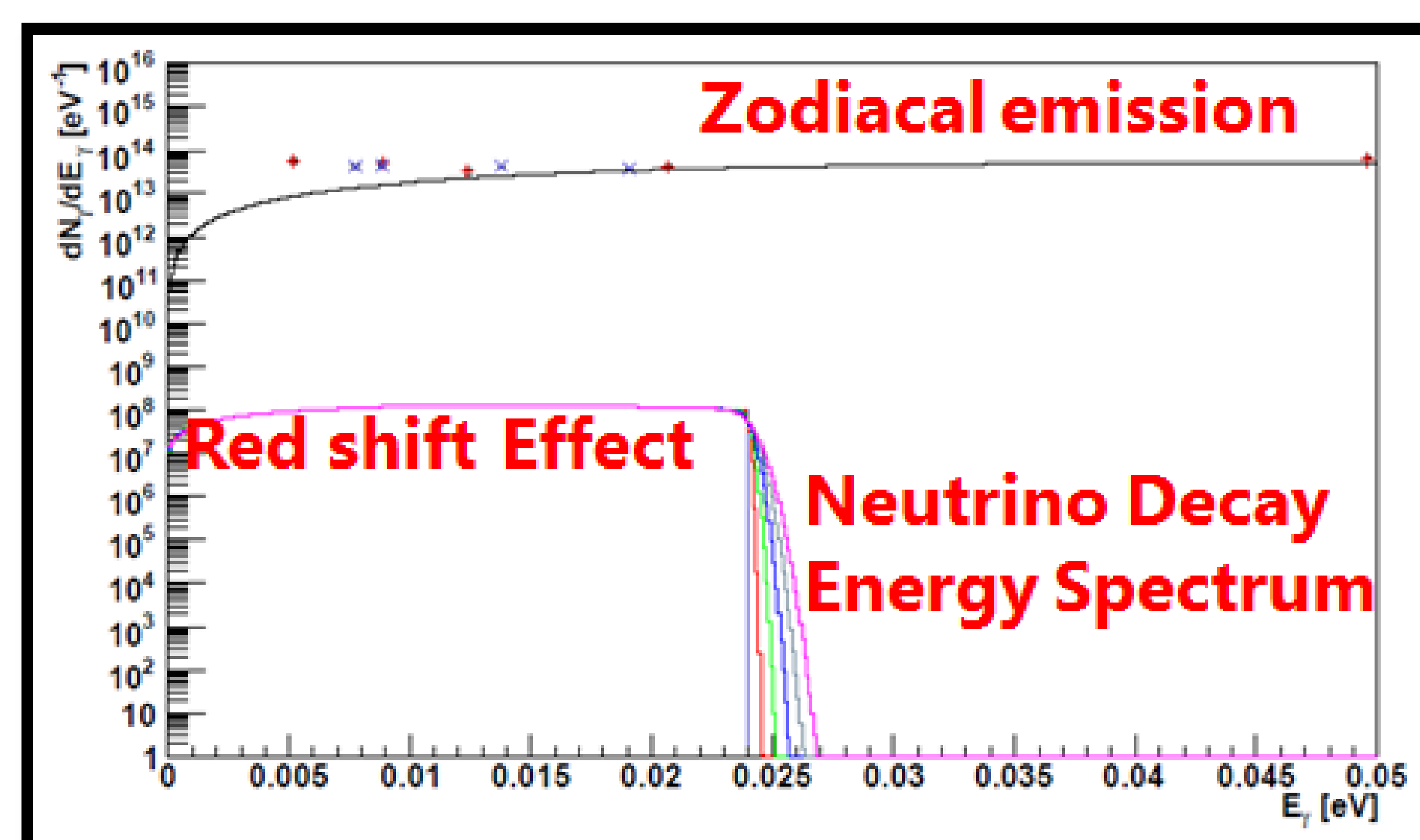
$$\begin{aligned} & \nu_3 \rightarrow \nu_2 + \gamma \\ & E_\gamma = \frac{\Delta m_{32}^2}{m_3} \end{aligned}$$



Feynman diagram of neutrino decay

- Neutrino lifetime is very long (10^{43} year in the standard model, 1.5×10^{17} year in Left-Right symmetric model).
- To observe neutrino decay, we need an immense quantity of neutrino.
- Most promising method is to observe the decay of the cosmic background neutrino(CBN).

- CBN has a particle density ρ of 110 cm^{-3} per generation



← Expected energy spectrum of CBN decay and background (Zodiacal emission)

- $m_3 = 50 \text{ meV}$ is assumed
- E_γ at neutrino restframe = 25 meV
- Energy spectrum of the photon from CBN has a cutoff at 25 meV and low energy tail due to red shift effect.

3. STJ detector

- **Detector requirements for CBN decay observation**
- Continuous spectrum of photon energy around $E_\gamma \sim 25 \text{ meV}$ ($\lambda = 50 \mu\text{m}$, far infrared photon).
- Energy measurement for single photon with better than 2% resolution for $E_\gamma \sim 25 \text{ meV}$ to identify the shape edge in the spectrum.

→ We adopted Hf-STJ as a detector for neutrino decay.

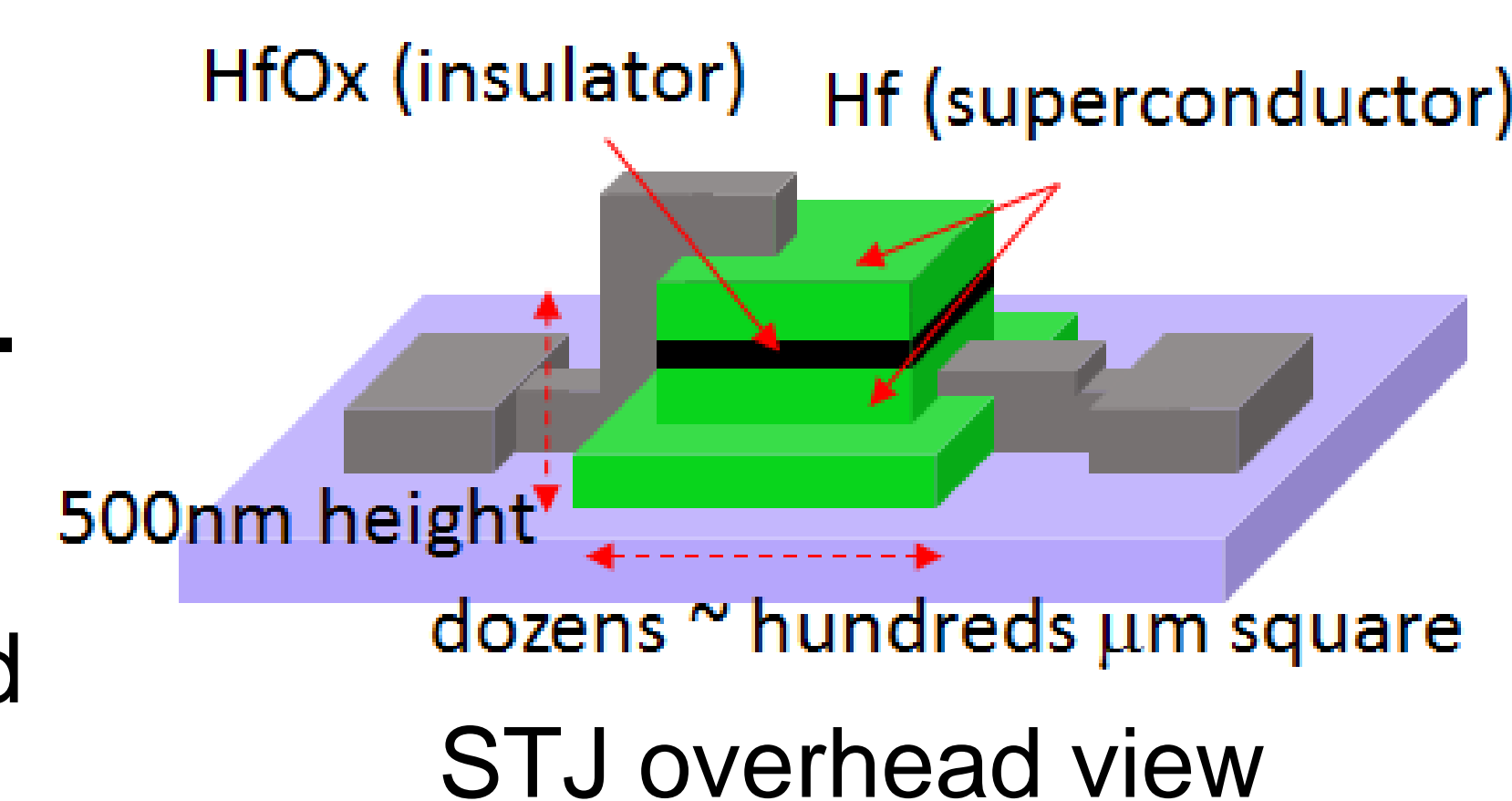
3. STJ detector

- **STJ = Superconductor / Insulator / Superconductor.**

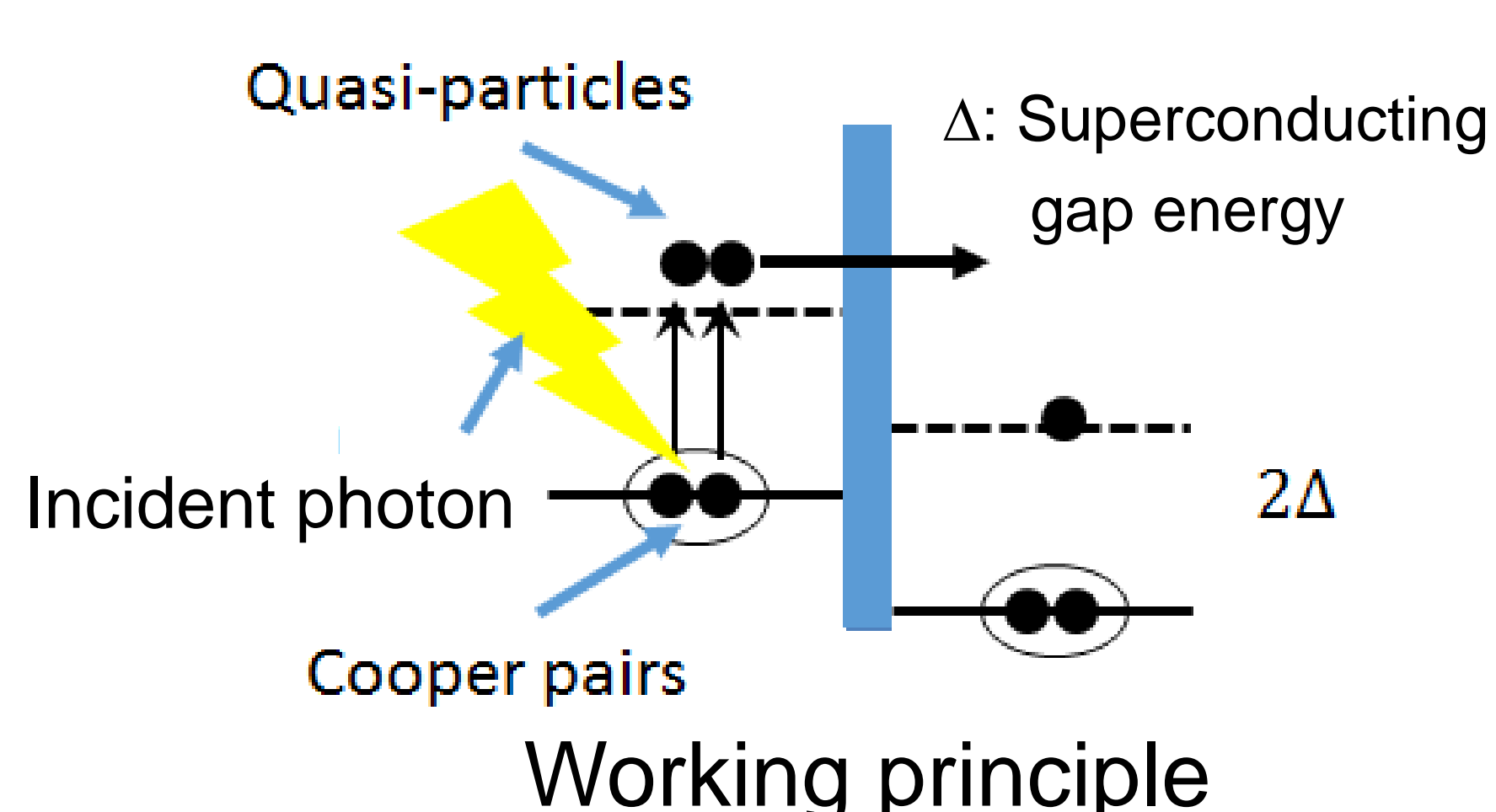
Characteristics:

- Single photon counting
- Broad response from the ultraviolet to the far infrared
- Arrival time information to $\sim 1 \mu\text{s}$
- High energy resolution

- **Must be operate at low temperature for best performance. ($T < 0.1 \times T_c$)**

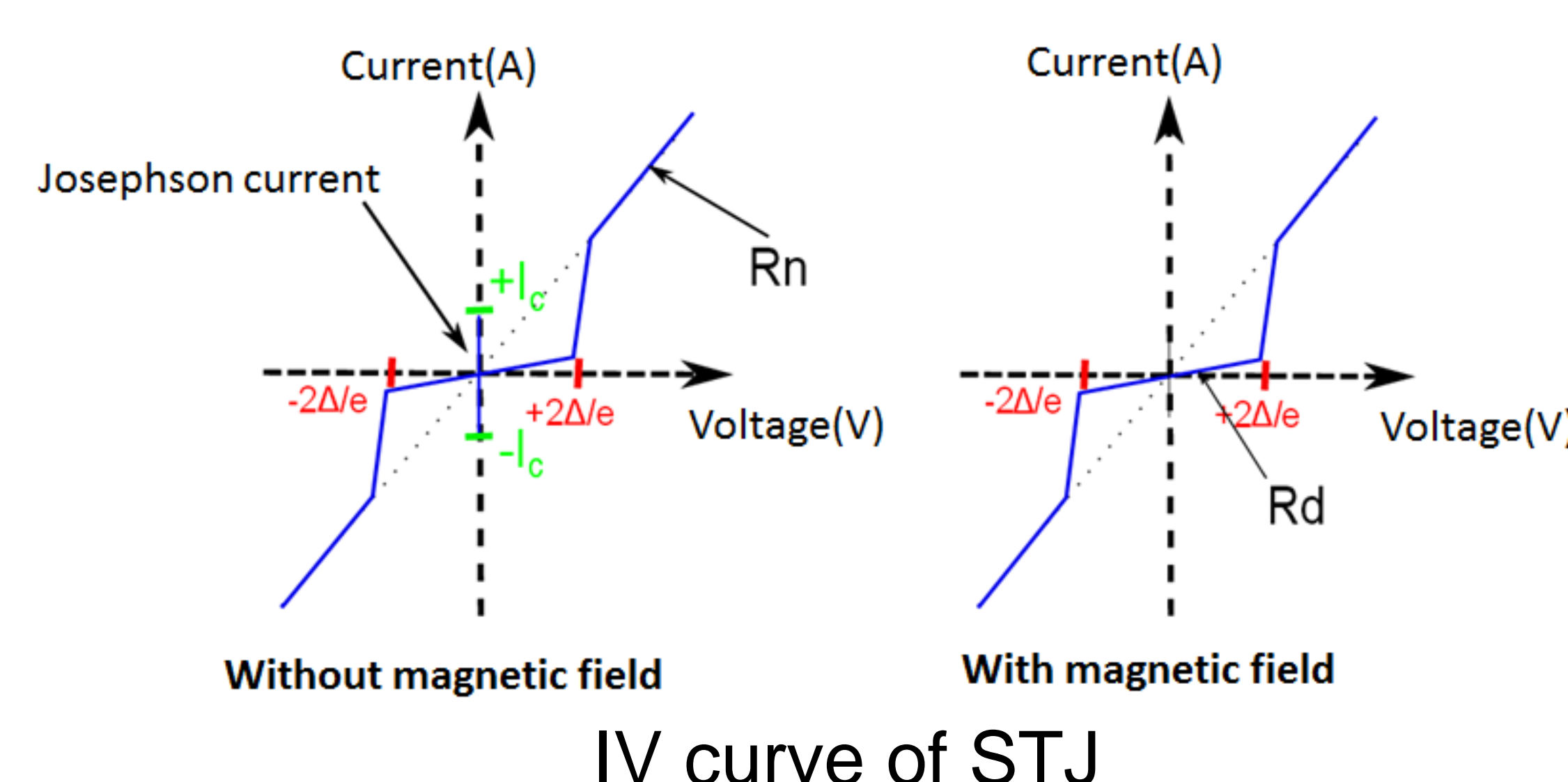


STJ overhead view



Working principle

1. Incident photon is absorbed in the superconductor and excites cooper pairs.
2. Exited cooper pairs become quasi-particles.
3. Quasi-particles go through tunnel barrier by tunnel effect.
4. Number of quasi-particle is determined by energy of incident particle.
5. Thus, we can measure the energy of incident particle by measuring the tunnel current.



IV curve of STJ

4. Energy resolution of STJ detector

- Statistical fluctuation in number of quasi-particle determines STJ energy resolution.
- Smaller superconducting gap energy Δ yields better energy resolution.

$$\sigma_E = \sqrt{(1.7\Delta)FE}$$

Material	Tc(K)	$\Delta(\text{meV})$
Niobium	9.20	1.550
Aluminum	1.14	0.172
Hafnium	0.13	0.021

Δ : Superconducting gap energy
F: fano factor
E: Photon energy

T_c : Superconducting critical temperature
Need $\sim 1/10 T_c$ for practical operation

Hf-STJ energy resolution

- $N_{\text{quasi-particle}} = 25 \text{ meV} / 1.7\Delta = 735$
- 2% energy resolution @ 25 meV is achievable if fano factor < 0.3

→ Hf-STJ can archive requirements for CBN decay observation.

5. Current status of development

- We observed Josephson current by Hf-HfOx-Hf barrier layer.
- So, we succeeded in producing Hf-STJ detector which works as STJ for the first time in the world.

Optimizations

- As our Hf-STJ sample has large leakage current, optimization is underway.

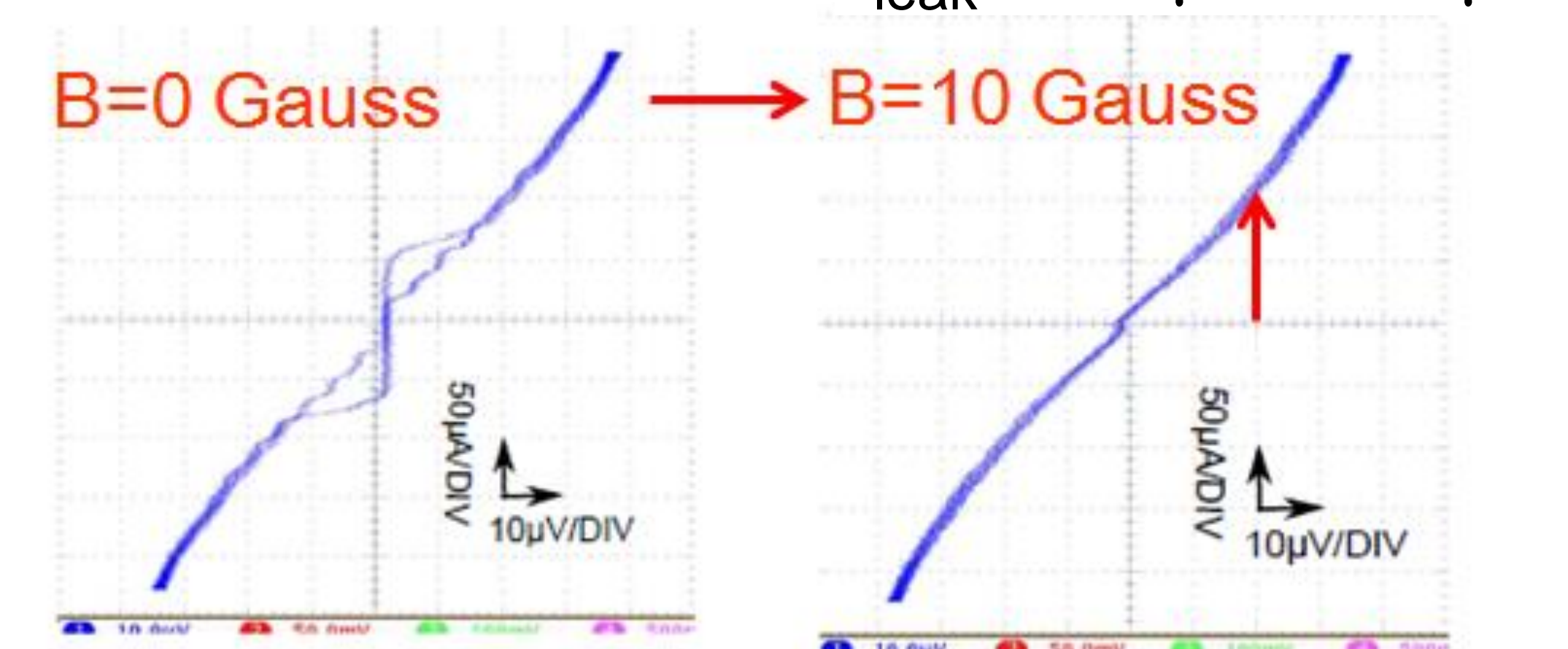
1. Downsizing

Hf-STJ($100 \times 100 \mu\text{m}^2$) shows smaller leakage current than Hf-STJ($200 \times 200 \mu\text{m}^2$).

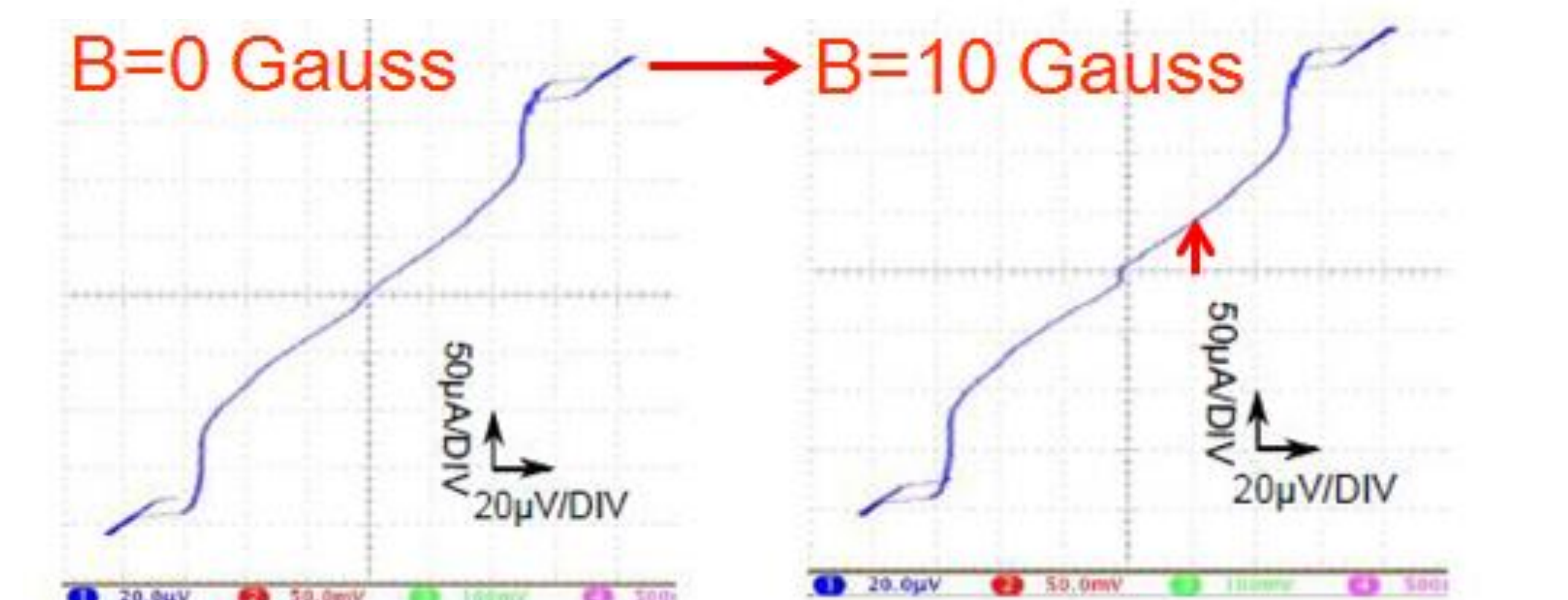
2. Optimize condition for making the insulator.

3. Oxidation on side surface.

I-V curve of Hf-STJ ($200 \times 200 \mu\text{m}^2$)
• $T \sim 80 \text{ mK}$, $I_c = 60 \mu\text{A}$, $R_d = 0.2 \Omega$

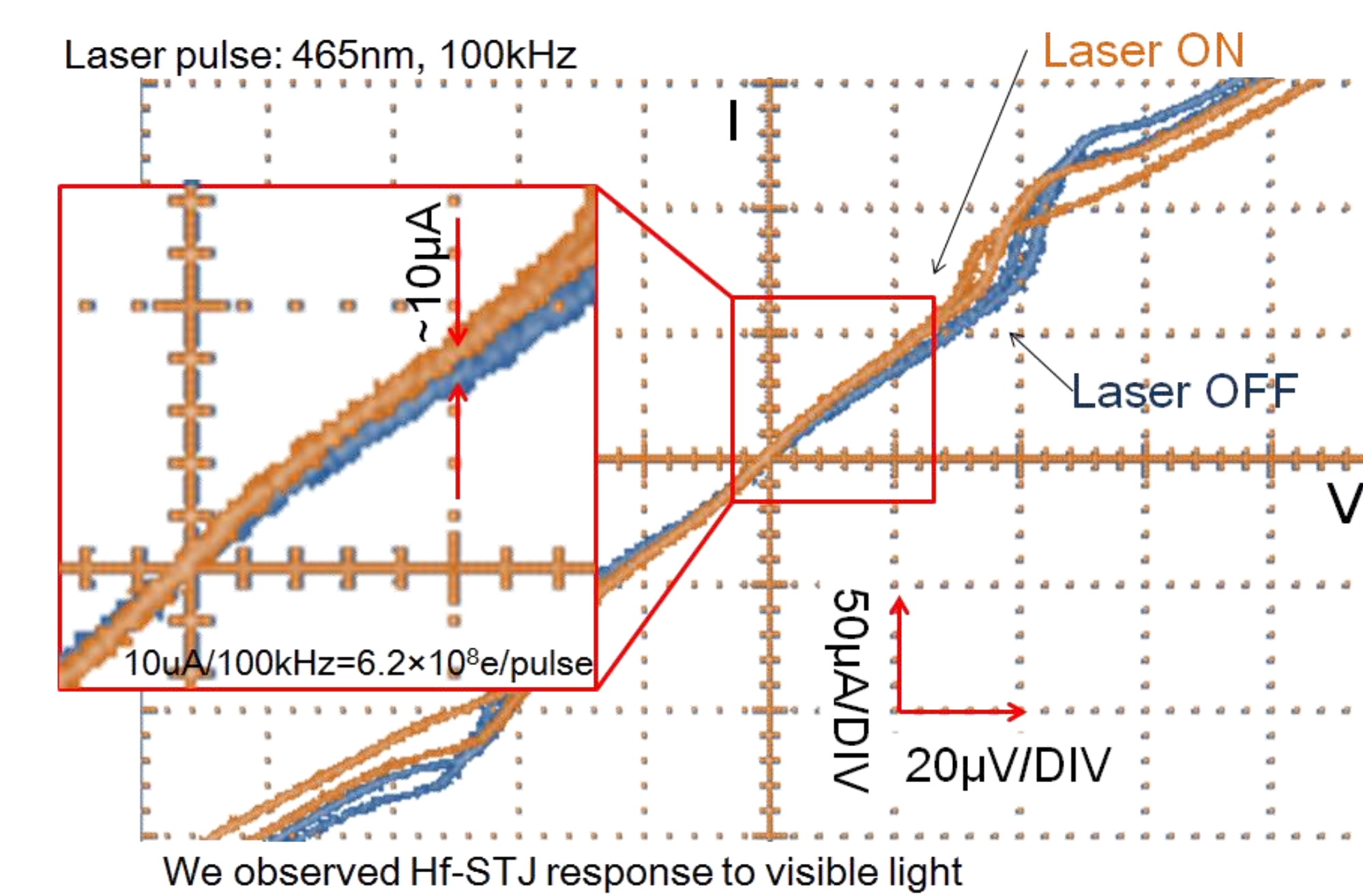


I-V curve of Hf-STJ ($100 \times 100 \mu\text{m}^2$)
• $T \sim 40 \text{ mK}$, $I_c = 10 \mu\text{A}$, $R_d = 0.6 \Omega$



Response to DC light

- We observed the response of Hf-STJ to DC-like laser light at wave length of 465 nm and frequency of 100 kHz .
- The signal current was measured to be $10 \mu\text{A}$ at $20 \mu\text{V}$.



← IV curve of Hf-STJ with and without 100kHz laser light(465 nm)

Sample information:
 $100 \times 100 \mu\text{m}^2$
 $R_d = 0.6 \Omega$
 $T = 139 \sim 153 \text{ mK}$

6. Summary

- To observe neutrino decay, we are developing Hf-STJ.
- We succeeded in producing Hf-STJ detectors which work as STJ for the first time in the world.
- Also, Hf-STJ response to the visible light was observed.
- However, our Hf-STJ sample has large leakage current of $30 \mu\text{A}$ @ $20 \mu\text{V}$ and need more improvement to function as a far-infrared photon detector.